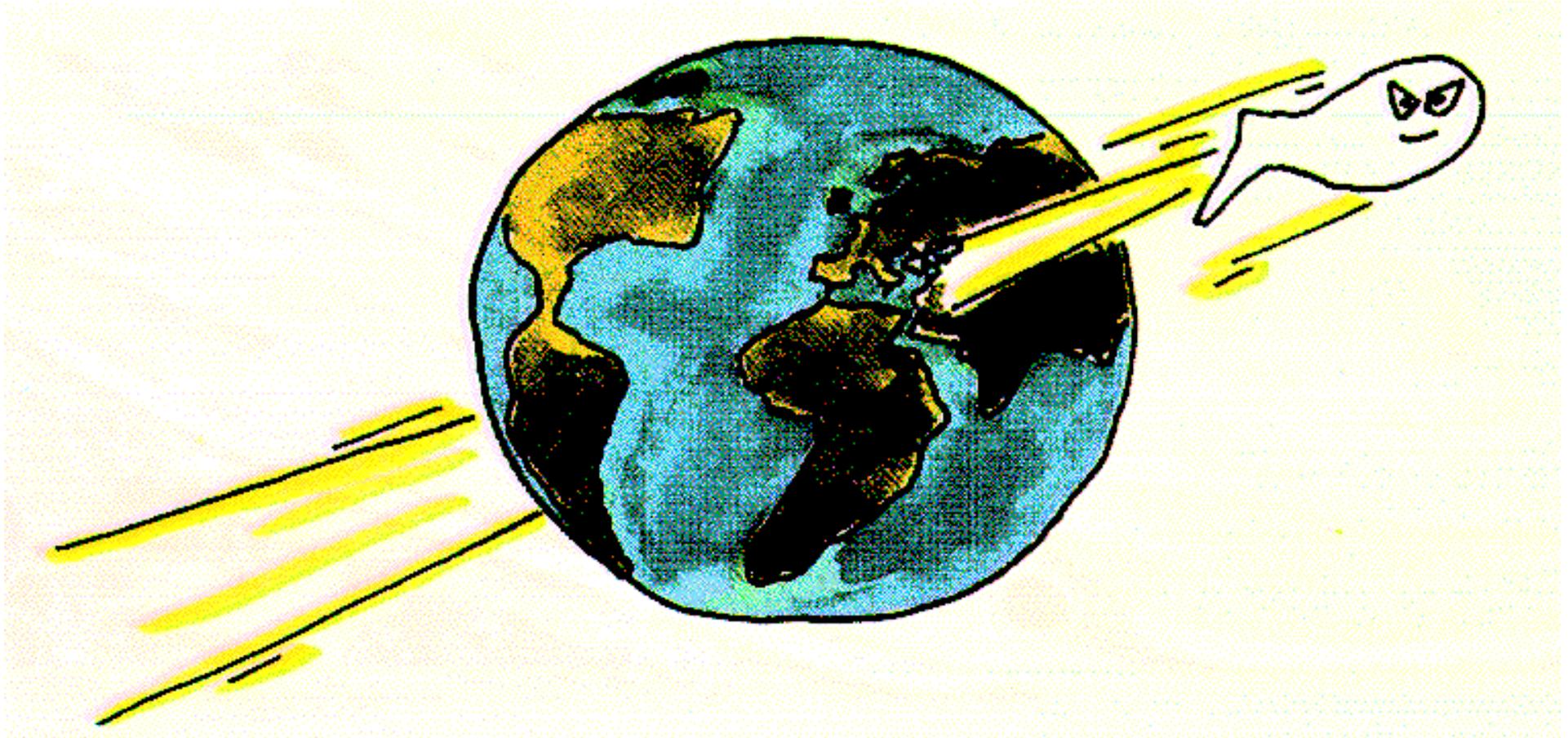


**Today:**

**Part I: Neutrino oscillations:  
beam experiments**

**Part II: Next tutorials:  
making distributions with  
histograms and ntuples**

# Super-Kamiokande Physics II: Long Baseline Beams



# Neutrino Oscillations

Assume

**FLAVOR STATES**

$$|\nu_f\rangle$$

weakly  
interacting

are  
superpositions  
of

**MASS STATES**

$$|\nu_m\rangle$$

unitary mixing matrix

$$|\nu_f\rangle = \sum_{i=1}^N U_{fi} |\nu_i\rangle$$

If mixing matrix is not diagonal,  
get *flavor oscillations*  
as neutrinos propagate

# Simple two-flavor

$$|\nu_f\rangle = \cos\theta |\nu_1\rangle + \sin\theta |\nu_2\rangle$$

$$|\nu_g\rangle = -\sin\theta |\nu_1\rangle + \cos\theta |\nu_2\rangle$$

Propagate a distance L:

$$|\nu_i(\mathbf{t})\rangle = e^{-iE_i t} |\nu_i(\mathbf{0})\rangle \sim e^{-im_i^2 L/2p} |\nu_i(\mathbf{0})\rangle$$

Probability of detecting flavor g at L:

$$P(\nu_f \rightarrow \nu_g) = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 L}{E} \right)$$

E in GeV  
L in km  
 $\Delta m^2$  in  $\text{eV}^2$

Parameters of nature to measure:  $\theta, \Delta m^2 = m_2^2 - m_1^2$

# The Experimental Game

- Start with some neutrinos (natural or artificial)
- Measure (or calculate) flavor composition and energy spectrum
- Let them propagate
- Measure flavor and energies again

Have the flavors and energies changed?

If so, does the change follow

$$P(\nu_f \rightarrow \nu_g) = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 L}{E} \right) \quad ?$$

Disappearance:  $\nu$ 's oscillate into 'invisible' flavor

e.g.  $\nu_e \rightarrow \nu_\mu$  at ~MeV energies



Appearance: directly see new flavor

e.g.  $\nu_\mu \rightarrow \nu_\tau$  at ~GeV energies



# Oscillation Parameter Space

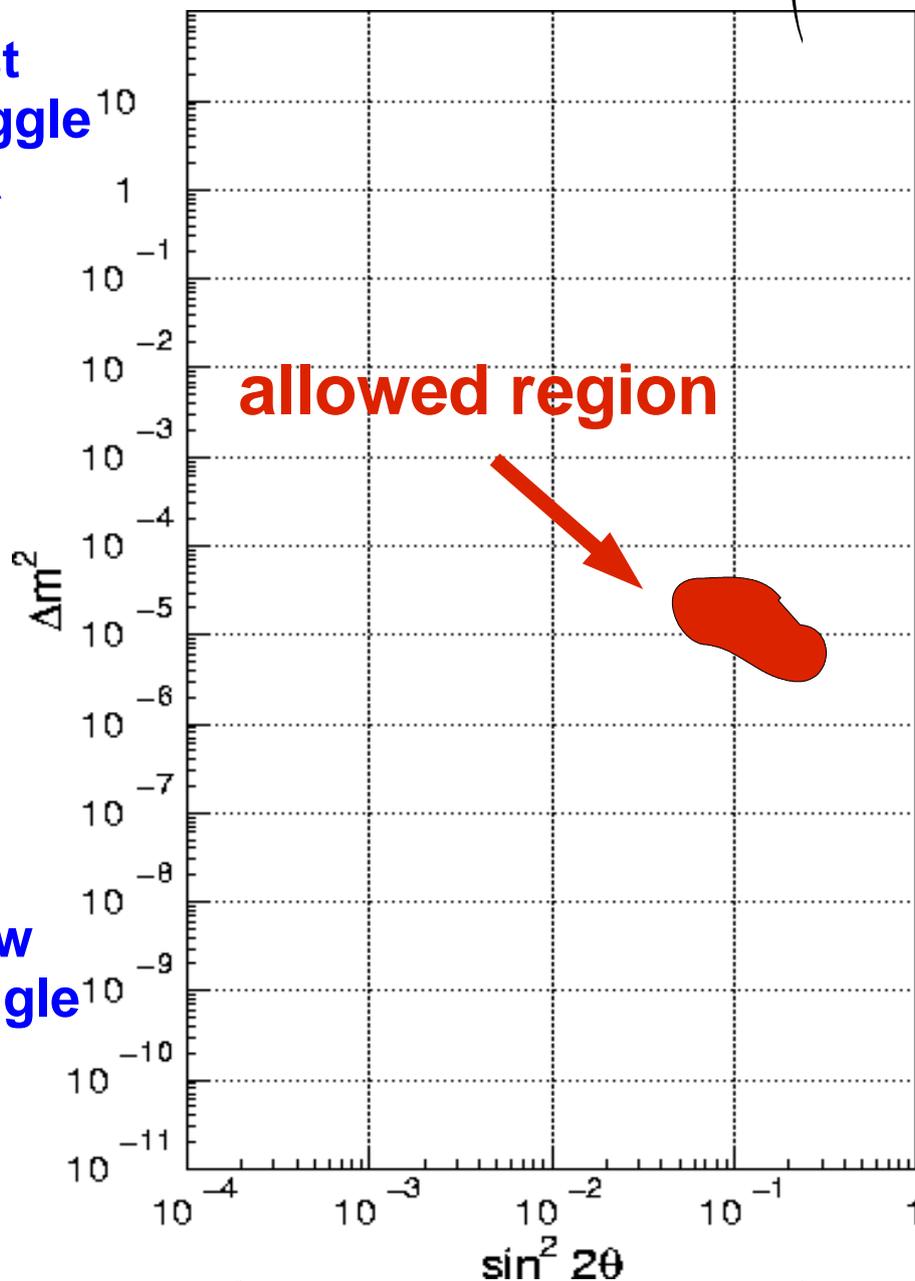
Twiddle  
L/E

Frequency  
 $\propto \Delta m^2 L/E$

fast  
wiggle

slow  
wiggle

$$P(\nu_f \rightarrow \nu_g) = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 L}{E} \right)$$

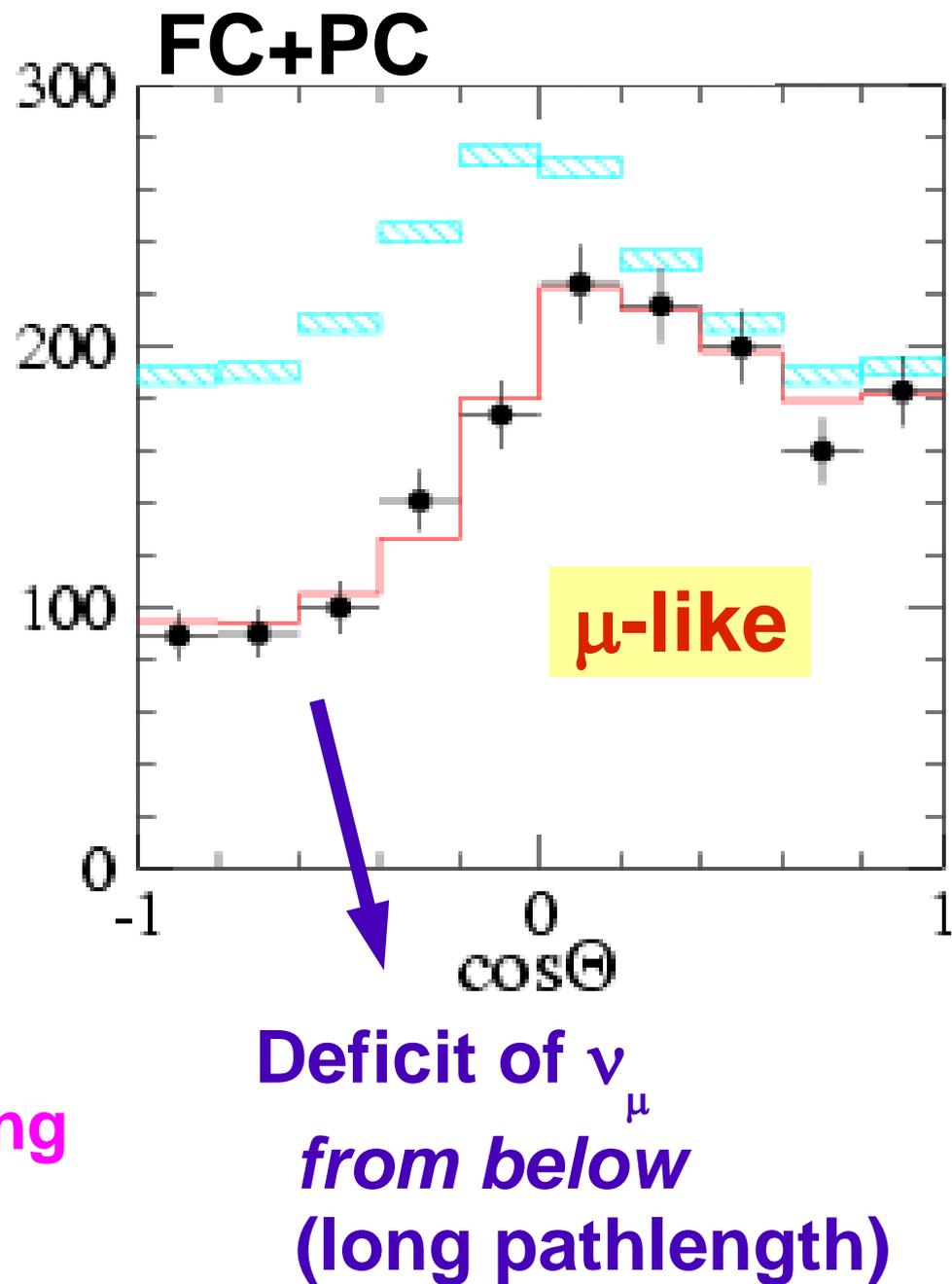
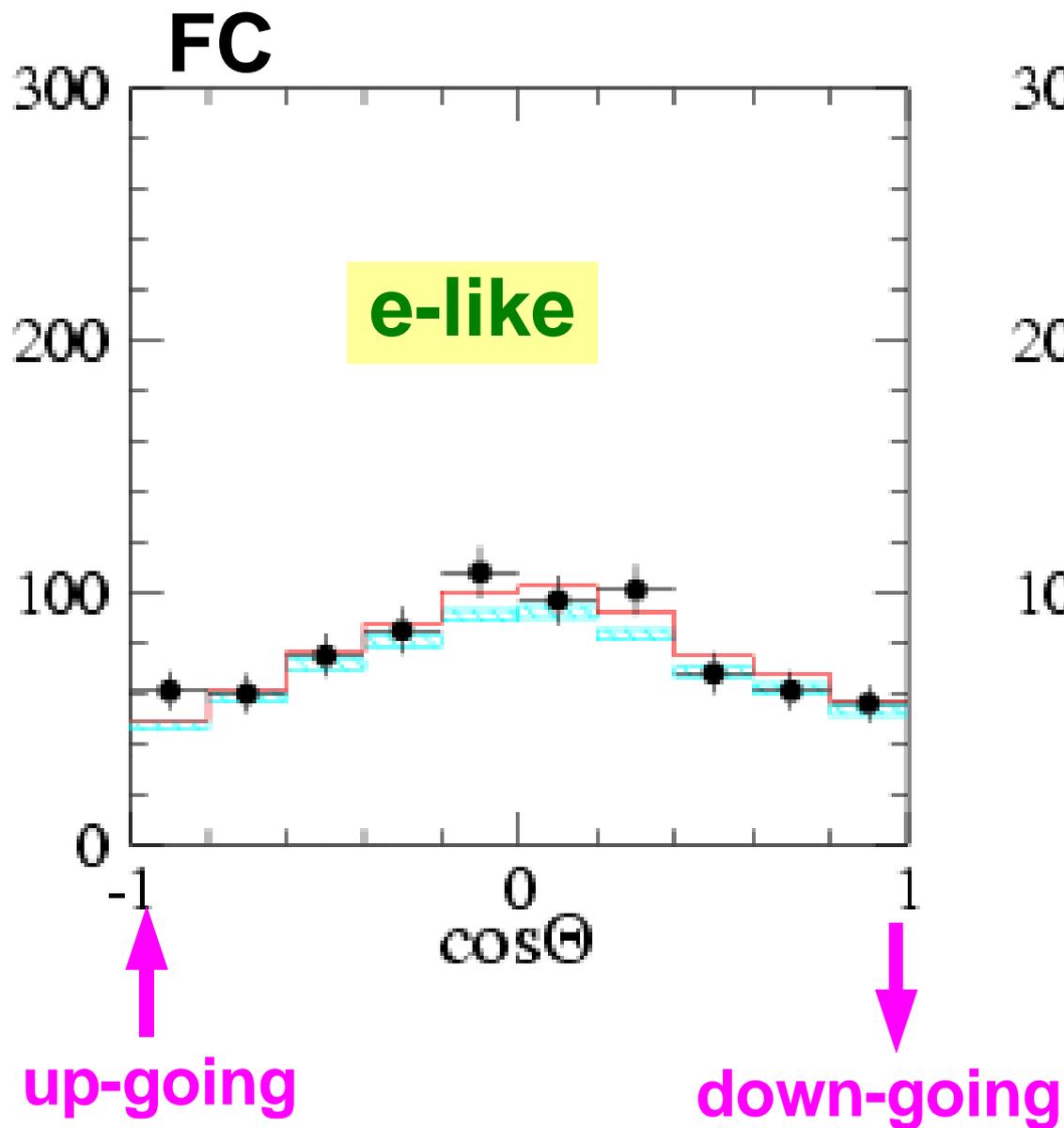


Amplitude  
 $\propto \sin^2 2\theta$

Experimental statistics

# Zenith angle distribution

1489 days of SK data (SKI)



**Next: INDEPENDENT TEST of atmospheric neutrino oscillations using a well-understood  $\nu$  beam**

**$E_\nu \sim \text{GeV}$ ,  $L \sim 100$ 's of km for same  $L/E$**

$$P(\nu_f \rightarrow \nu_g) = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 L}{E} \right)$$

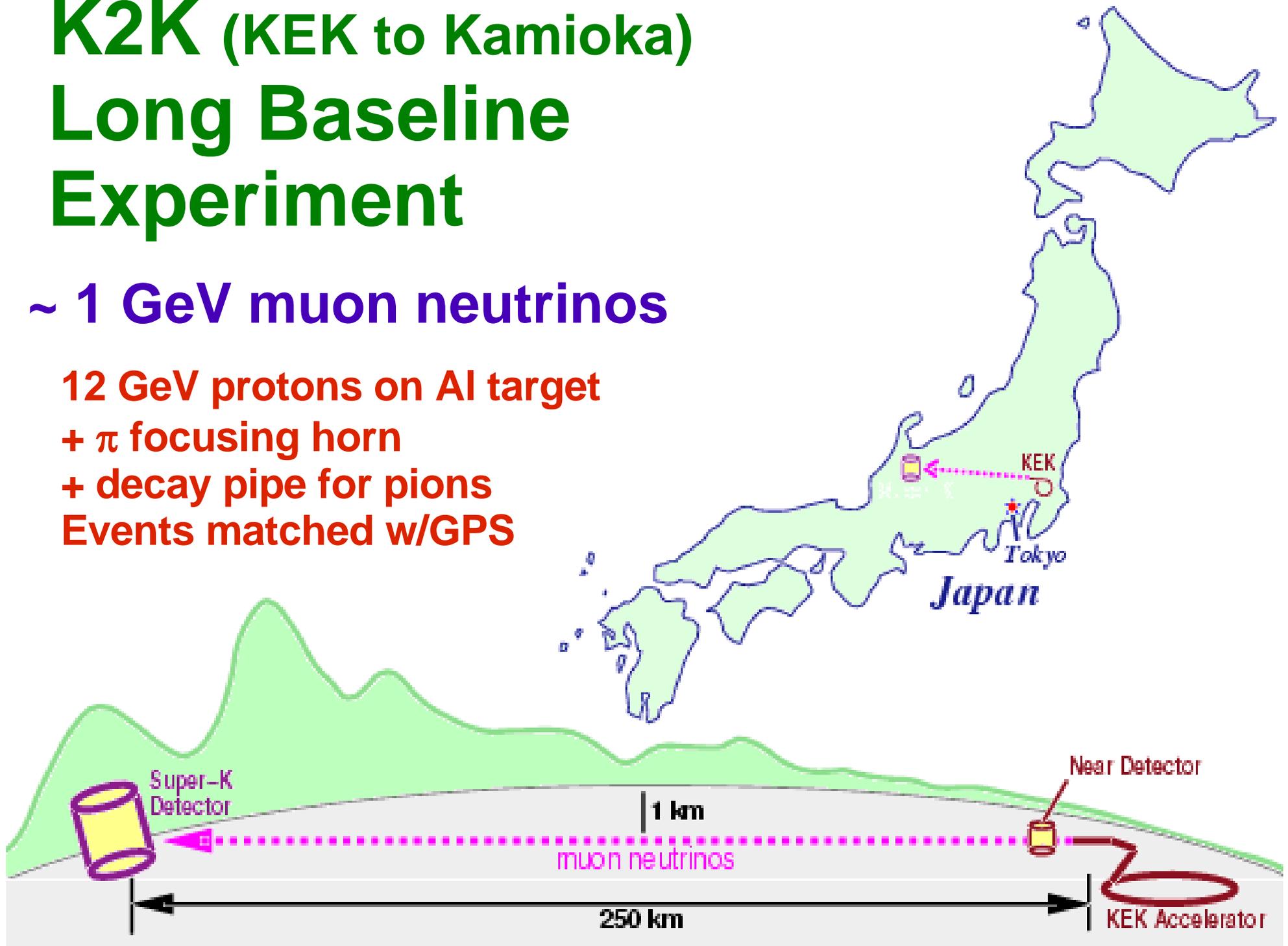
## **LONG BASELINE EXPERIMENTS**

**Compare flux, flavor and energy spectrum at near and far detectors**

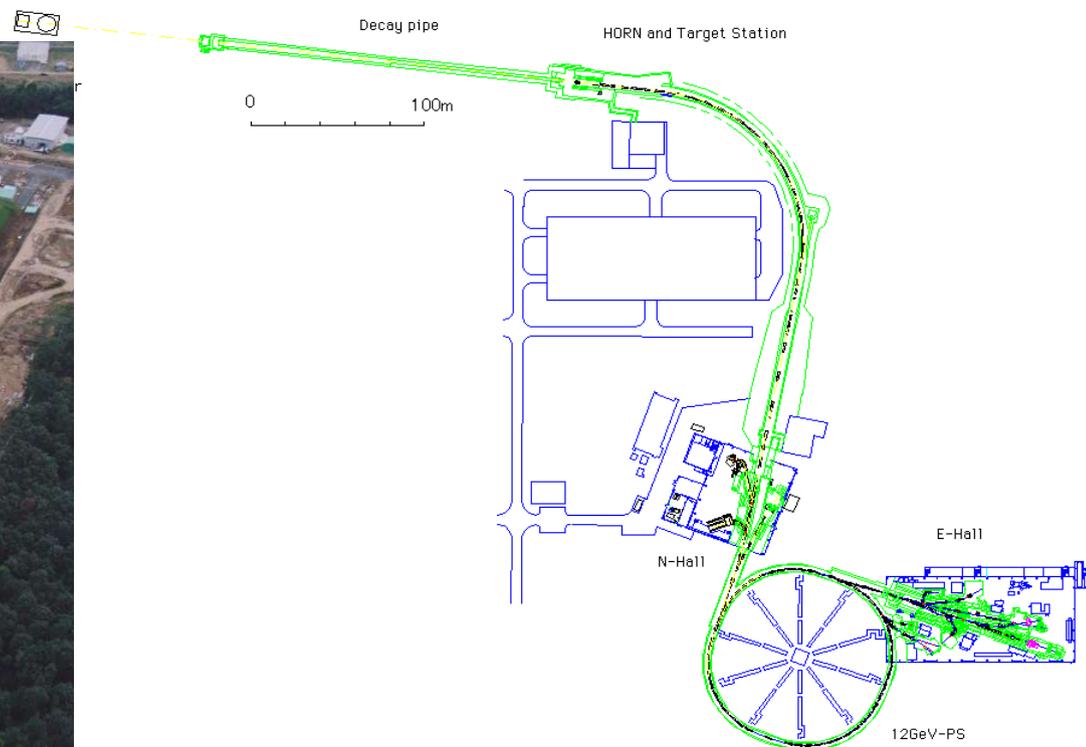
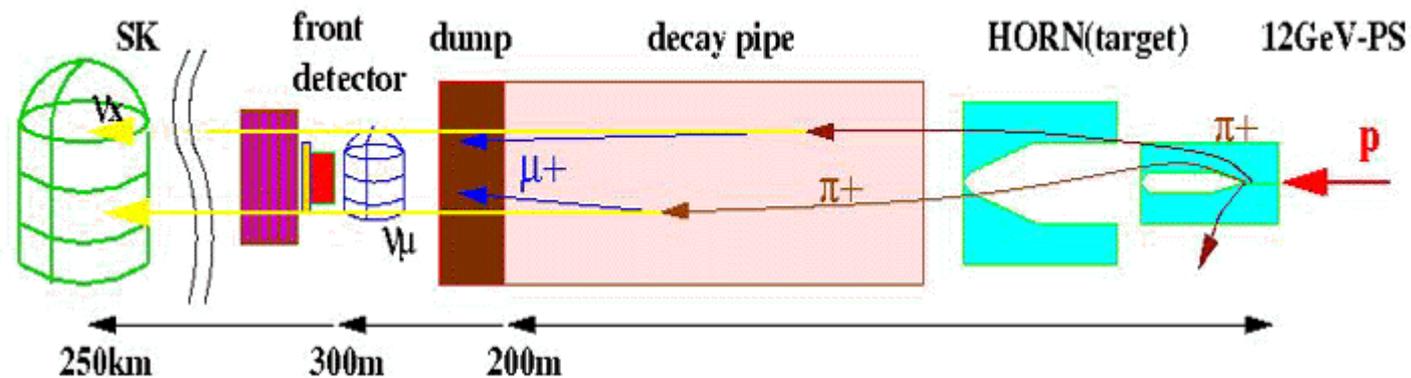
# K2K (KEK to Kamioka) Long Baseline Experiment

~ 1 GeV muon neutrinos

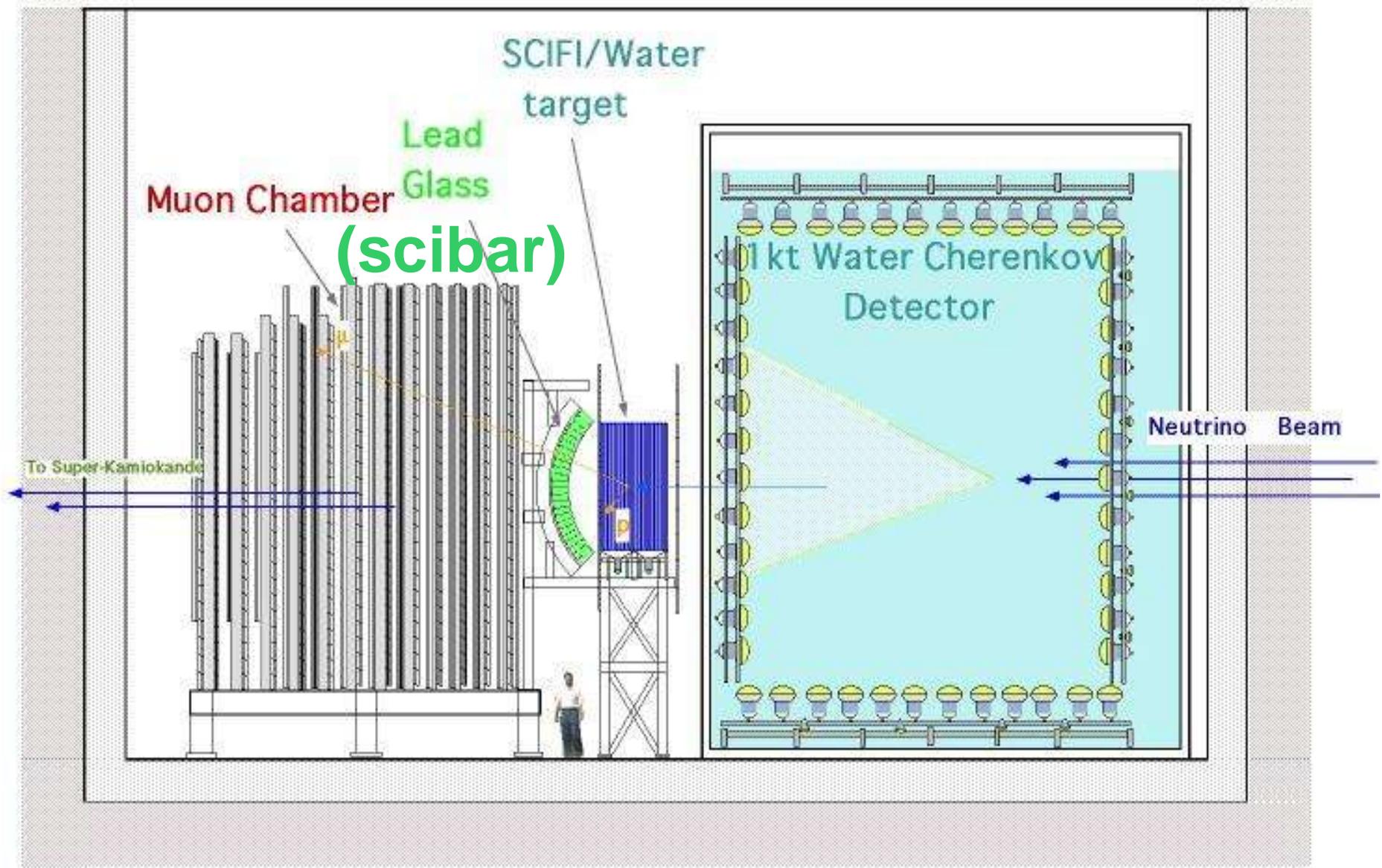
12 GeV protons on Al target  
+  $\pi$  focusing horn  
+ decay pipe for pions  
Events matched w/GPS



# The Neutrino Beamline at KEK

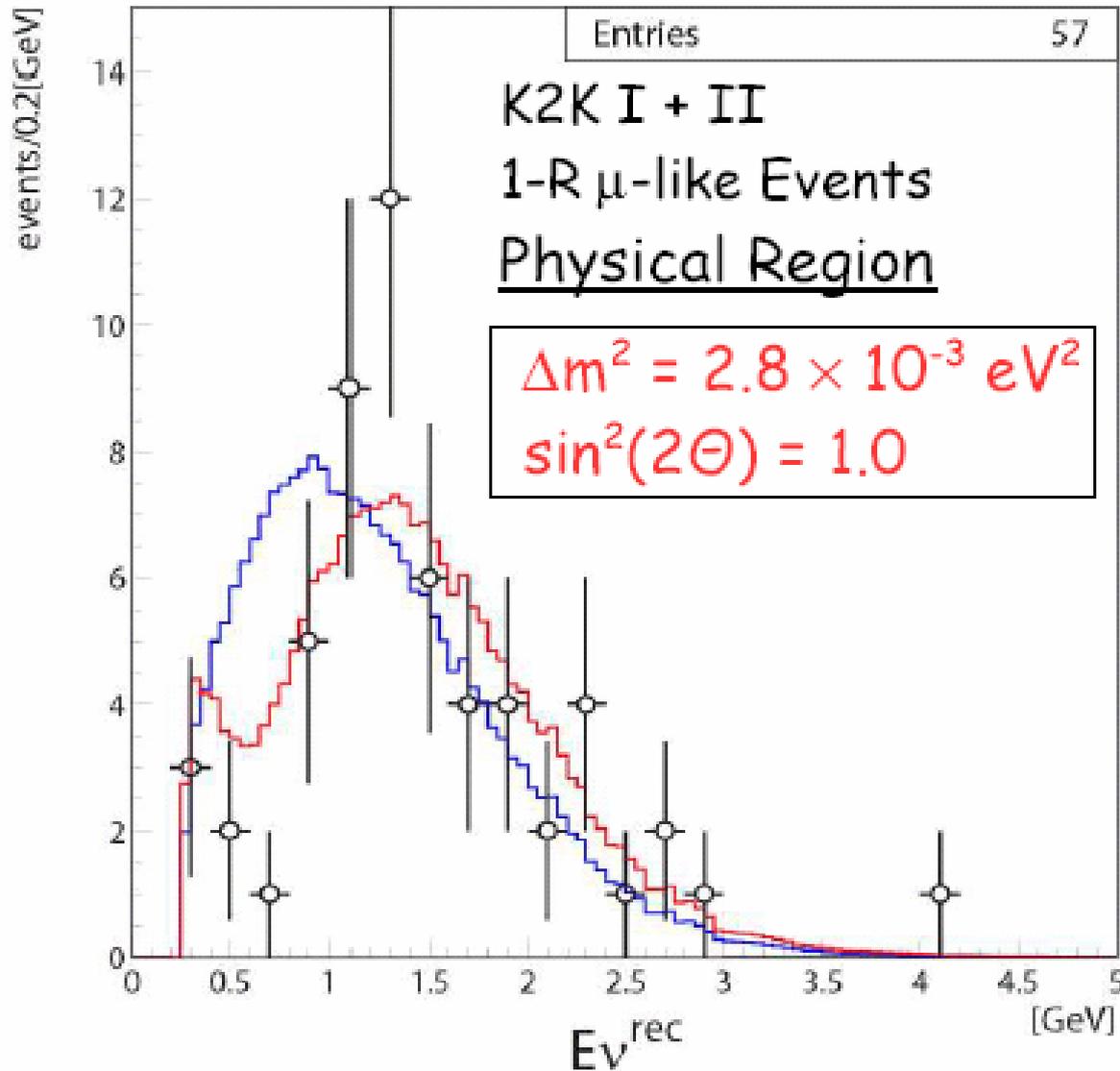


# The Near Detector (300 m away)



**Characterize the  $\nu$  beam for extrapolation to SK**

# Results from K2K

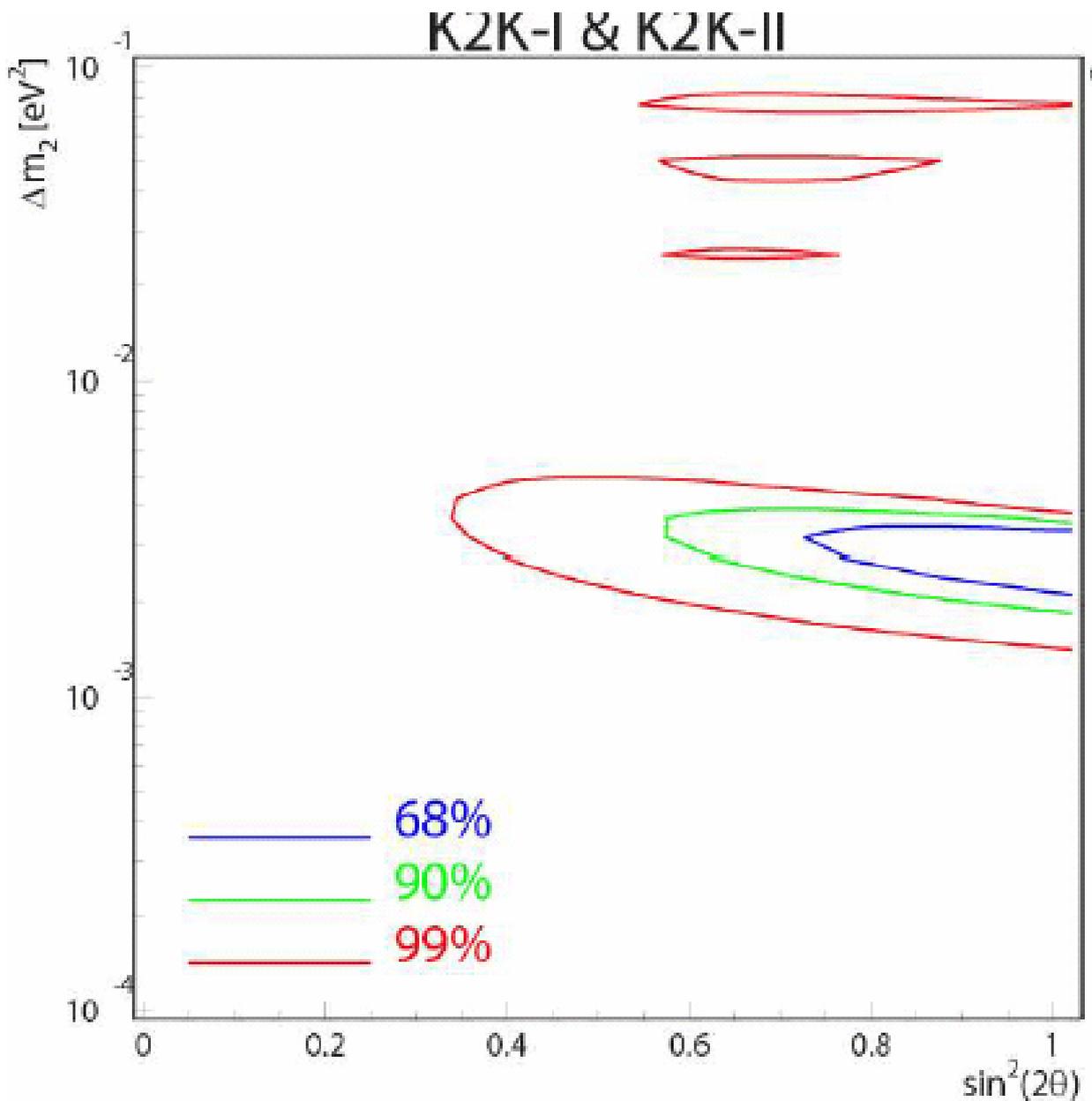


Single-ring  
 $\mu$ -like  
events

Total 107  
beam events  
observed;  
expect 149.7

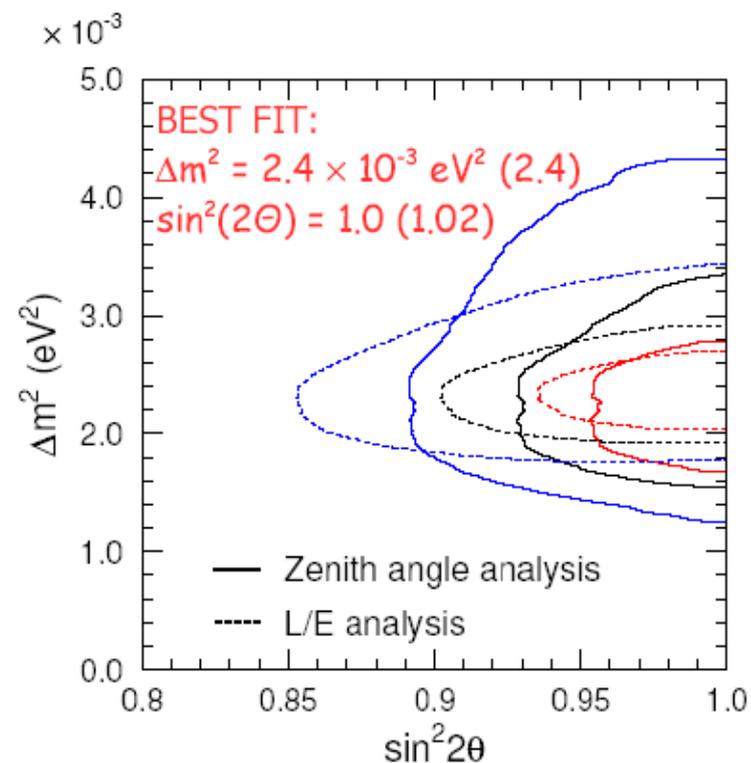
Suppression observed,  
spectral distortion consistent with oscillations

# K2K Allowed Oscillation Parameters



**Best Fit Results:**  
 $\Delta m^2 = 2.8 \times 10^{-3} \text{ eV}^2$   
 $\sin^2(2\theta) = 1.0$   
(constrained to physical region)

**Consistent with SK atmospheric**



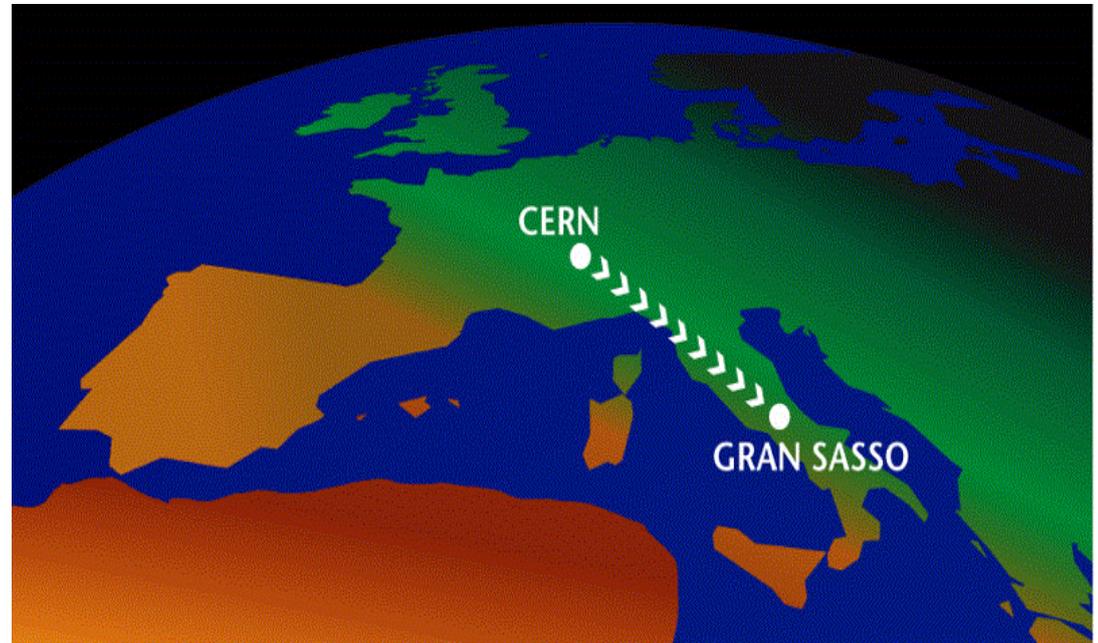
# Future Long Baseline Experiments

start 2005-6, 730 km distance

MINOS: Fermilab  
to Soudan



CNGS: CERN to  
Gran Sasso

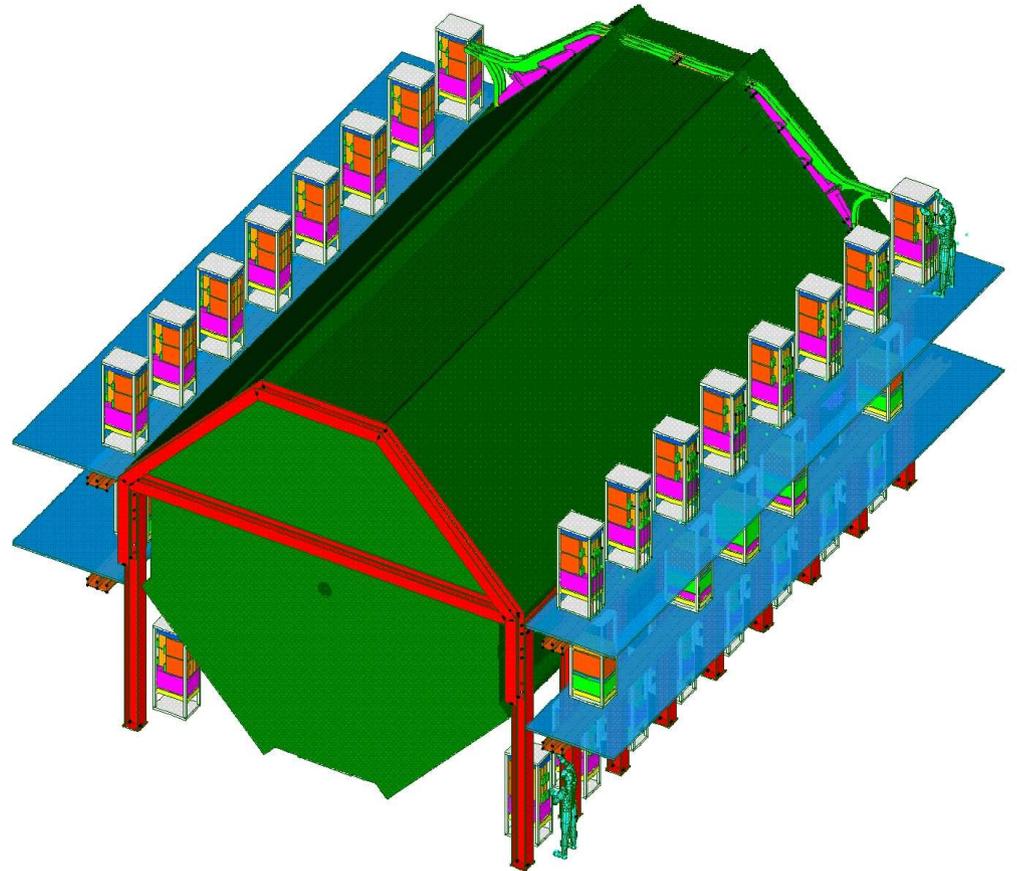


# NuMI Beamline at Fermilab



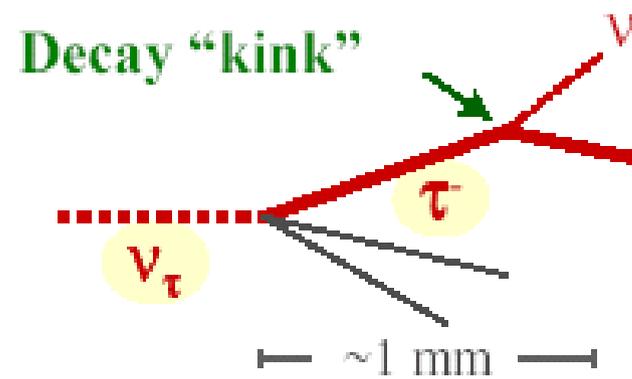
## MINOS Detector at the Soudan mine

iron plates  
+ scintillating fibers  
w/ magnetic field

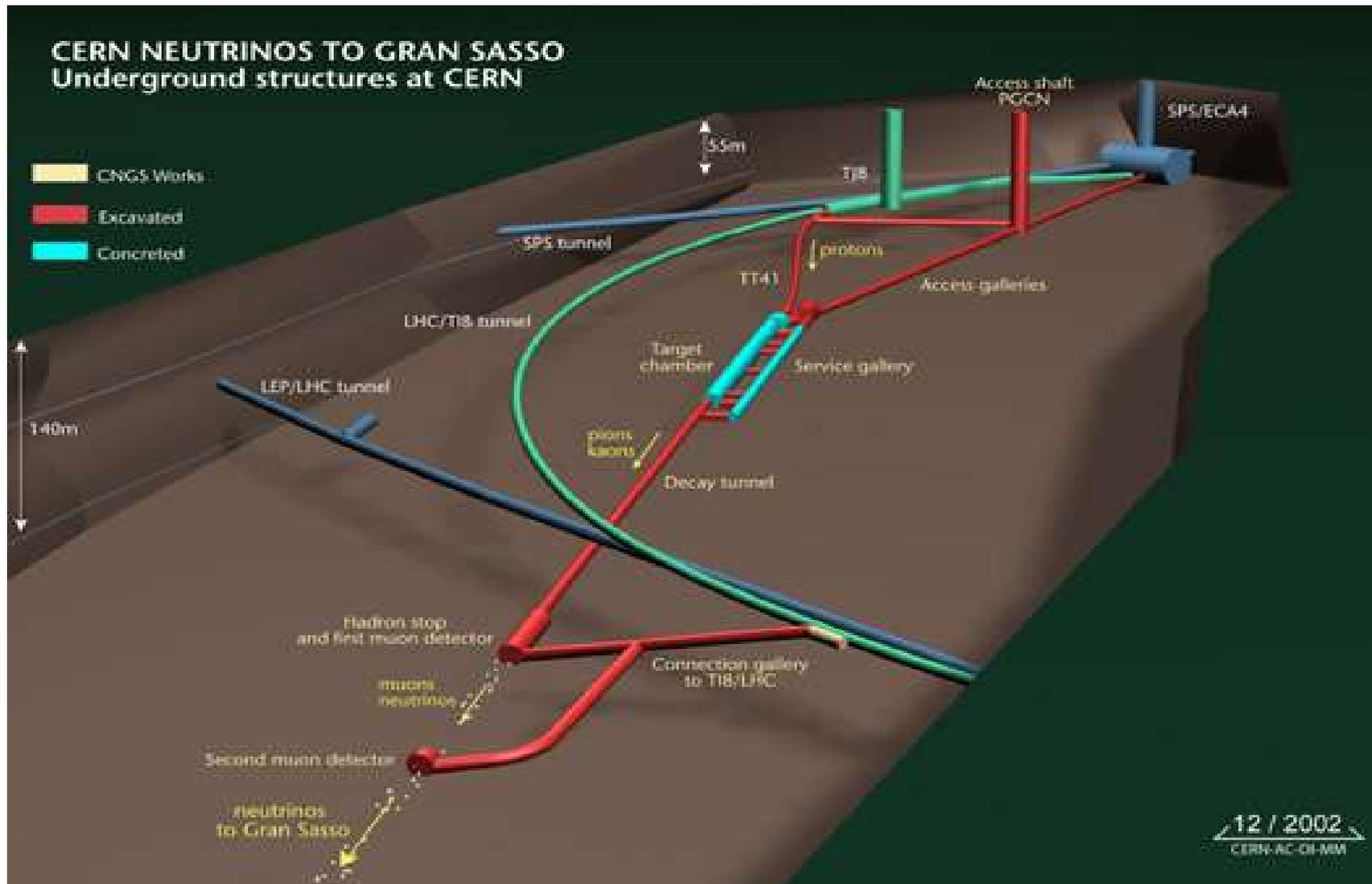


# CNGS

## CERN Neutrinos to Gran Sasso

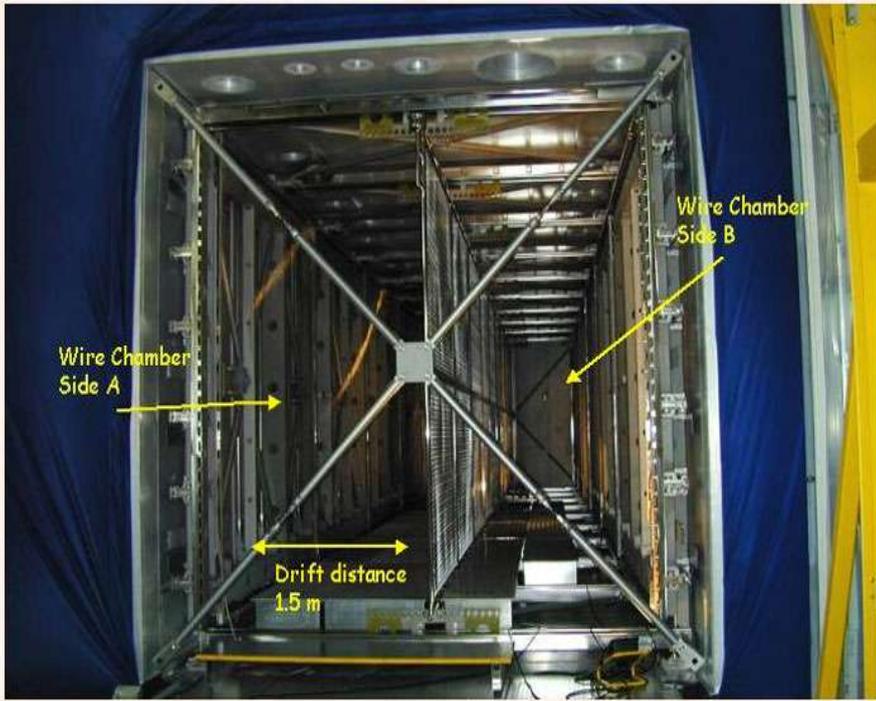


Look for  
 $\tau$  neutrinos  
explicitly



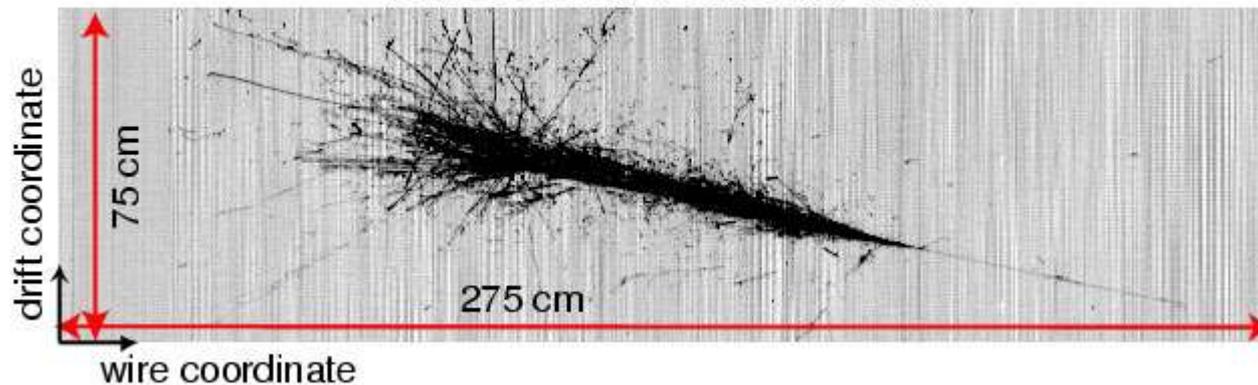
# ICARUS

## Liquid Argon Time Projection Chamber



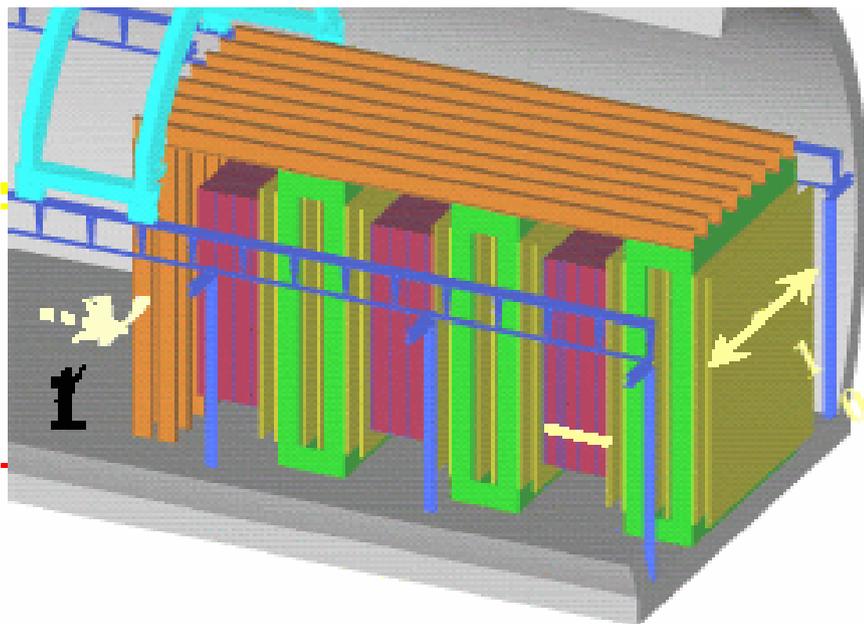
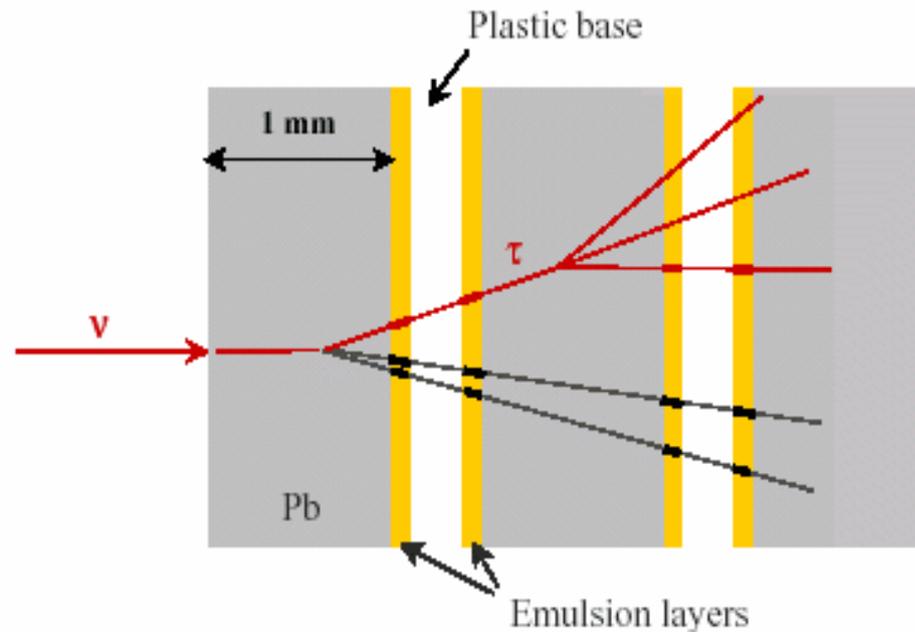
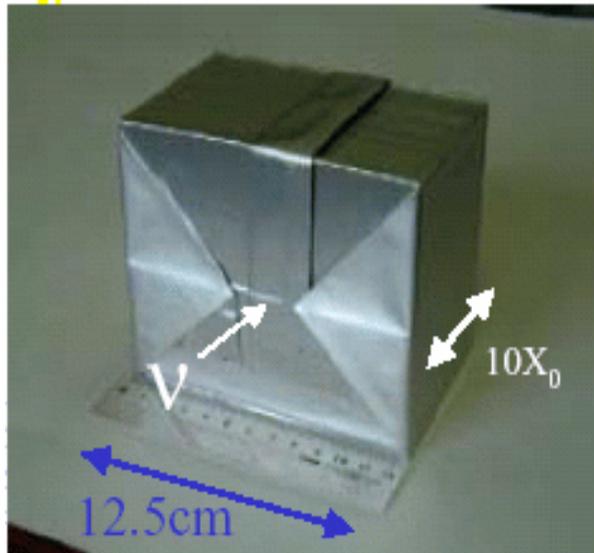
"Digital Bubble Chamber"

Run 308 Event 7 Collection view



# OPERA

lead/emulsion sandwich +  
scintillator planes



Extract bricks for  
scanning if  
electronic detector  
indicates  $\tau$ -like event

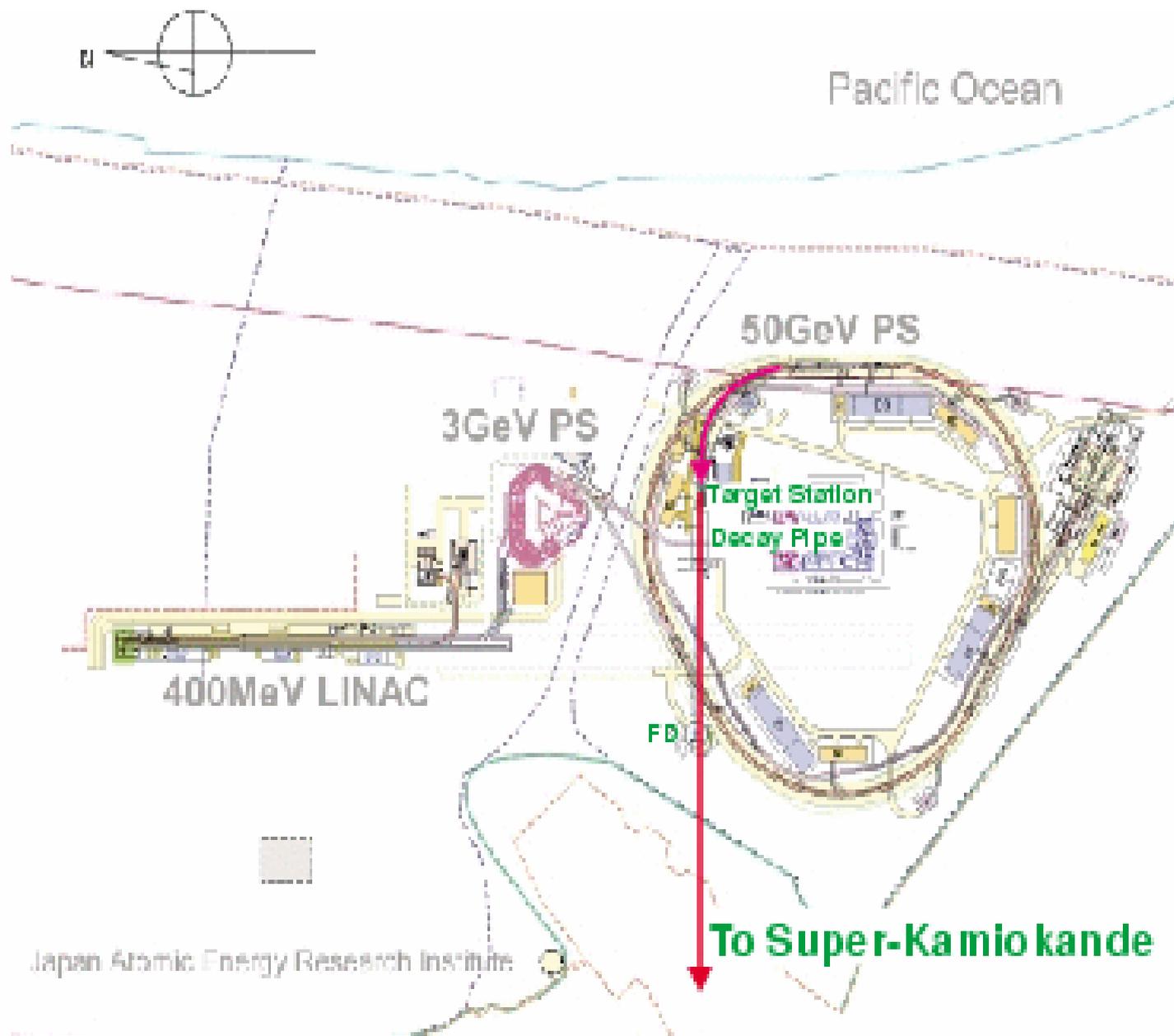
# T2K: "Tokai to Kamioka"



**Existing detector: Super-K**  
295 km, about 30 times K2K flux  
2.5 deg. off axis

# T2K ("Tokai to Kamioka")

2° off-axis

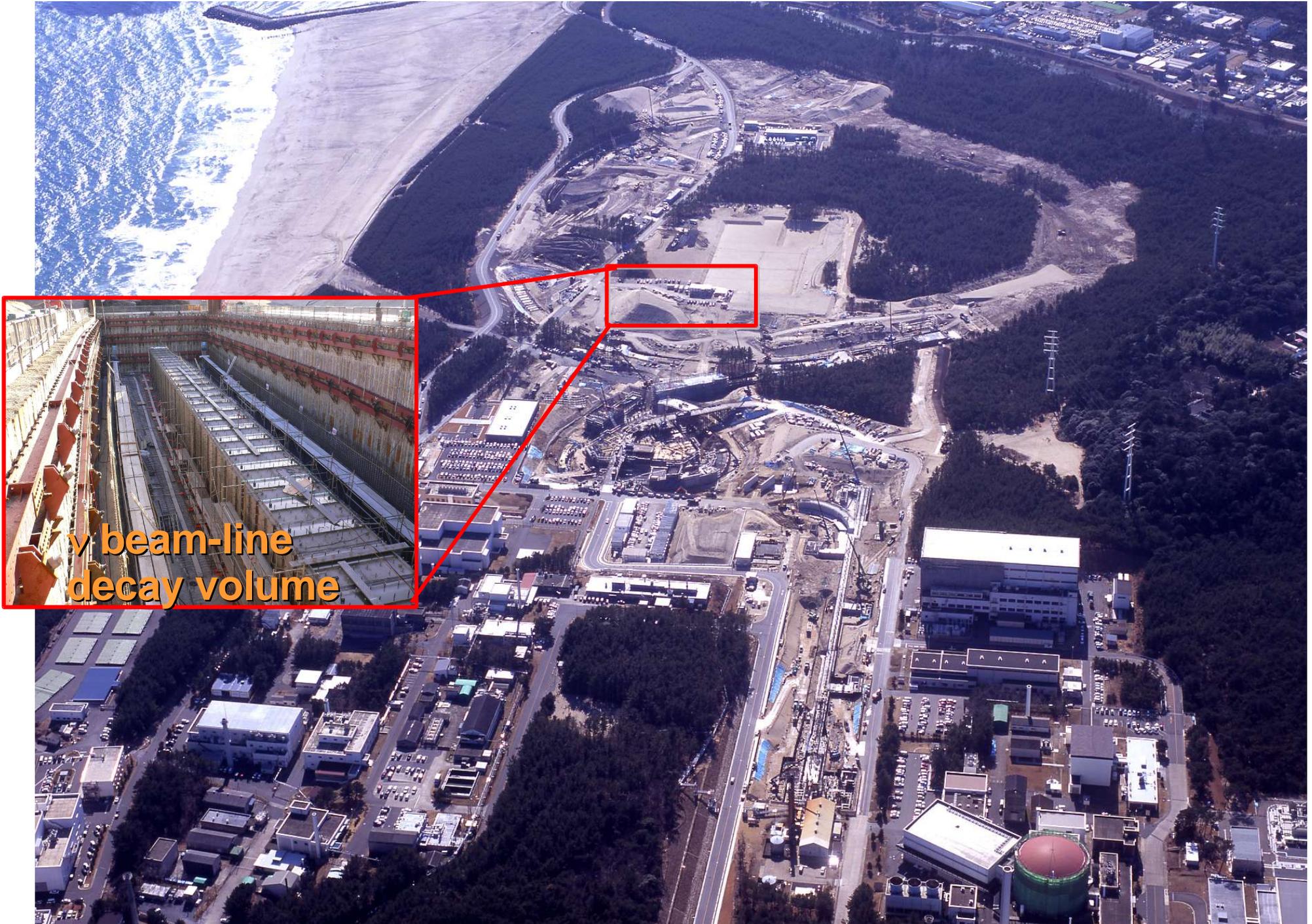


Near  
+ 2 km  
detectors,  
Super-K  
refurbished  
by 2006

Start 2009

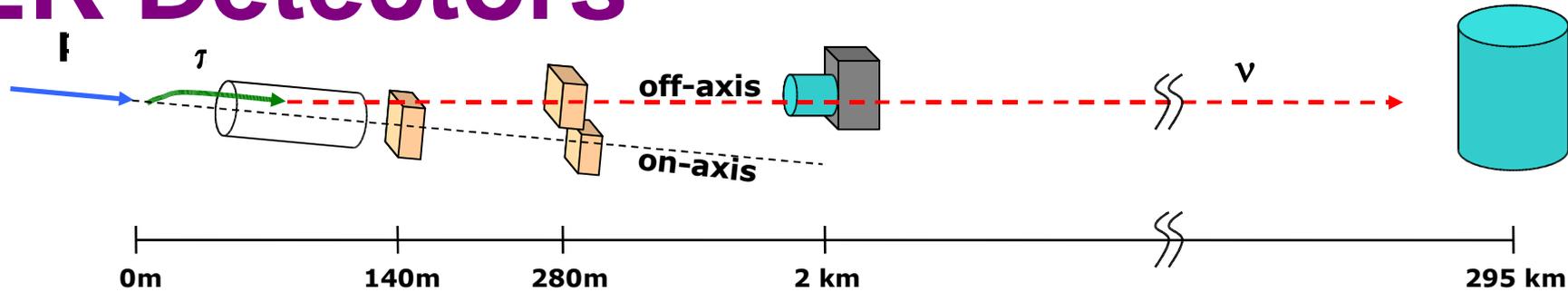
Beam  
neutrino  
energy  
tuned to  
oscillation  
minimum

# J-PARC beamline under construction now



gamma beam-line  
decay volume

# T2K Detectors

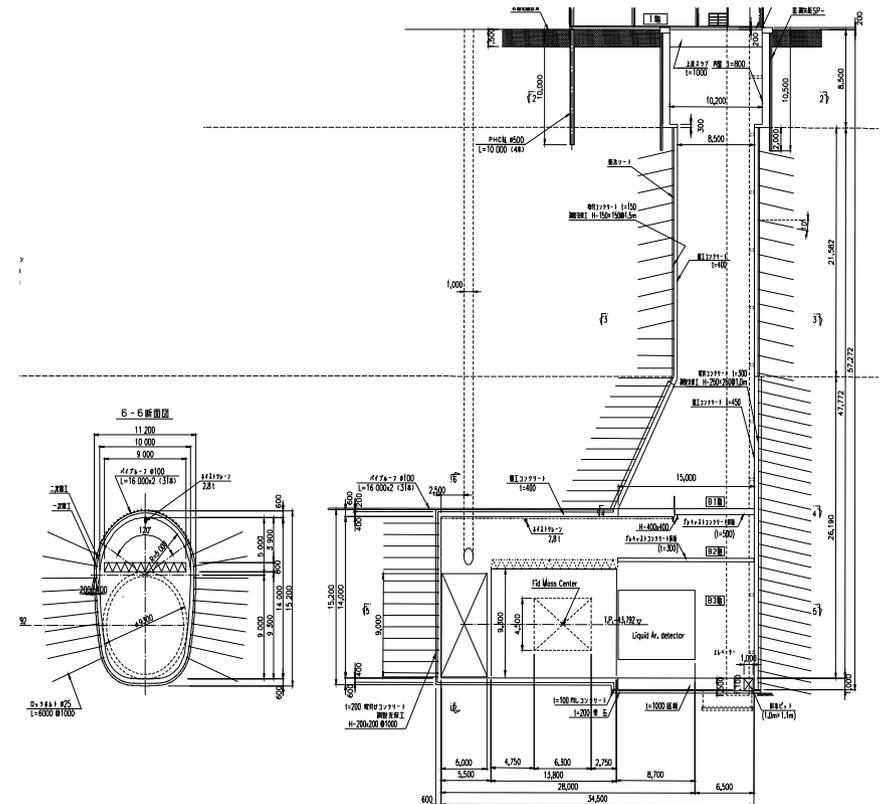


1. Near detector at 280 m: fine-grained tracker

2. Intermediate detector at 2 km

Water Cherenkov + fine-grain tracker (LAr) + muon ranger

Beam spectrum looks more like beam at SK  
⇒ *cancel systematics*



3. Super-K III at 295 km, fully refurbished

# Summary

**Atmospheric muon neutrinos are changing flavor ('disappearing') as they travel through the Earth:**

**This measurement in SK was the first clear indication of neutrino mass and oscillation**

$$P(\nu_f \rightarrow \nu_g) = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 L}{E} \right)$$

**This oscillation disappearance has now been confirmed with a long-baseline beam of neutrinos: K2K results are consistent**

**Next generation: MINOS, CNGS, T2K**

# SK Data Tutorial Today

1. Run superscan to eyeball data files



2.1 Set up and run sample program to look at charges and times for events

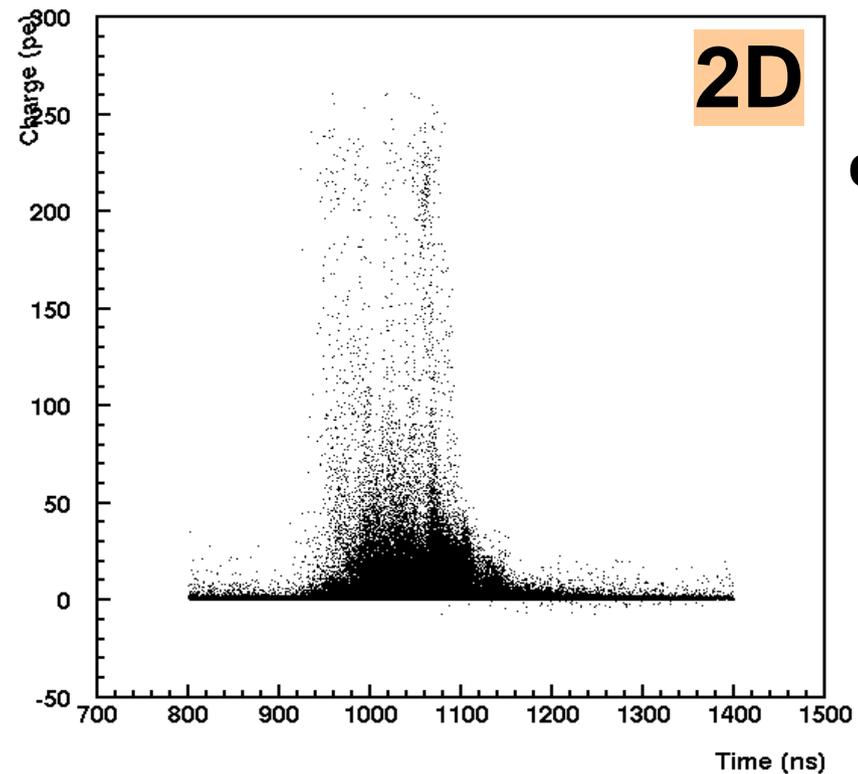
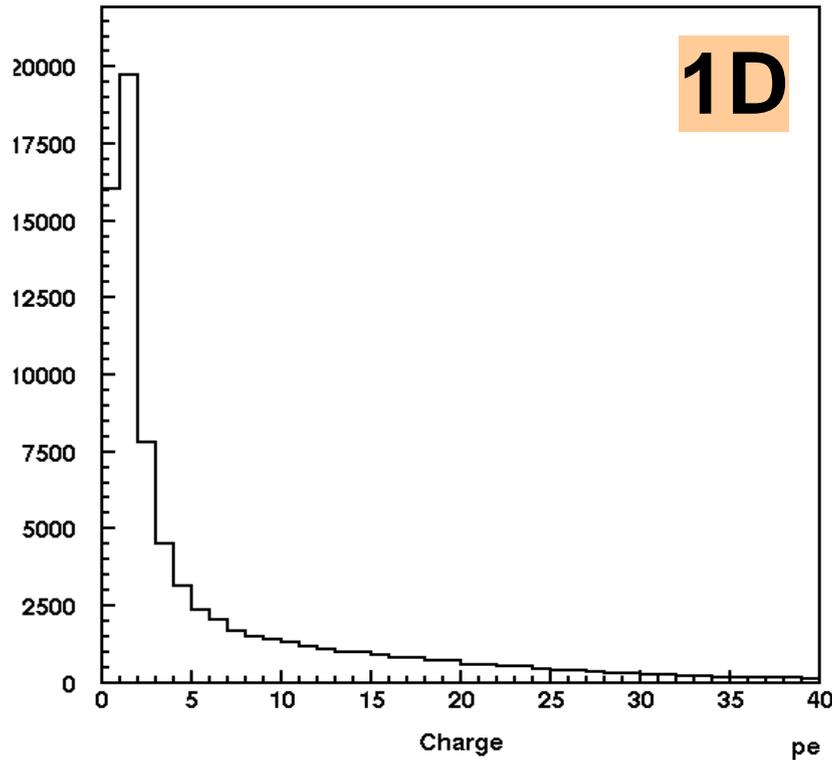


2.2 Make and view histograms (distributions)

**TODAY**

- Histograms made directly in program
- Ntuples made using PAW

# Histogram: a distribution of some variable(s)



density  
of dots  
 $\propto$  number  
of events  
in bin

**Cernlib/PAW: software package that allows manipulation of histogram structures**

- create histogram
- 'fill' with data
- output histo to file

# Ntuple: list of 'events' and their variables

e.g. run no., event no. , total charge, mean t

2000, 1, 42.,1020.

2000, 2, 13.,982.

2000, 3, 22.,984.

Note: an 'event' need not be an SK event (trigger);  
you can make an tuple of tube hits, or any  
kinds of items ('event'  $\equiv$  an item in the list)

**PAW will allow you to plot distributions of  
any chosen variable, or any variable  
vs any other, with 'cuts' (conditions)**

e.g. `nt/plot 1.q%t t>800&&t<1000`

means plot charge vs time for times  
between 800 and 1000

**very  
powerful!**

## Making a histogram in the program:

- Open data file (skopenf)
- Loop over events
  - Read one event (skread)  
(information stashed locally)
  - Loop over PMT hits
    - For each hit, fill charge  
and time histograms
- Close data file (skclosef)
- Output histogram to file

Later, view histogram with PAW (or Root)

## Making simple ntuples:

- Open data file (skopenf)
- Loop over events
  - Read one event (skread)  
(information stashed locally)
  - Loop over PMT hits
    - For each hit, output list of variables to a text file
- Close data file (skclosef)

Later, use PAW to make ntuples from text file,  
and plot desired variables

**Next 2 weeks: I will be in Japan, but other HEP profs (who would like to recruit students for research) will speak**